

# Curvature-Dependent CP Asymmetry and Gravitational Selection in Baryogenesis

Russell Crawford

**Abstract:** *The observed matter-antimatter asymmetry of the universe, characterized by a baryon-to-photon ratio  $\eta_B \approx 6 \times 10^{-10}$ , is commonly attributed to CP-violating amplitudes in local quantum field interactions treated against a nearly flat spacetime background. This study examines the matter-antimatter asymmetry of the universe, characterized by a baryon-to-photon ratio  $\eta_B \approx 6 \times 10^{-10}$ . We propose a framework in which CP violation, spacetime curvature, and global energy balance are dynamically coupled in the early universe. Using an effective field theory approach in curved spacetime, we show that CP-violating amplitudes can receive curvature-dependent corrections that become dominant at Planck-scale curvature, yielding  $\epsilon_{CP}$  values up to  $10^{-4}$  to  $10^{-8}$ . We further demonstrate that the Hamiltonian constraint, combined with photon binding energy, restricts viable baryogenesis scenarios and naturally selects small matter excess consistent with observations. Additionally, nonlinear curvature effects enhance the localization of photon energy into matter over antimatter, amplifying asymmetry. Order-of-magnitude estimates indicate that the observed  $\eta_B$  can arise for coefficients  $c_1 \approx 10^{-2}$ . The framework predicts observable signatures in gravitational waves and cosmic microwave background non-Gaussianity, suggesting a gravitational contribution to baryogenesis.*

**Keywords:** baryogenesis; CP violation; curved spacetime; effective field theory; gravitational energy; matter-antimatter asymmetry; Misner-Sharp energy

## 1. Introduction

The universe is overwhelmingly dominated by matter; antimatter appears only in trace amounts. Standard baryogenesis attributes this asymmetry to small CP-violating amplitudes in particle interactions under the three Sakharov conditions: (i) baryon number violation, (ii) C and CP violation, and (iii) departure from thermal equilibrium [1]. In most treatments these processes are analyzed in effectively flat spacetime, with gravity as a passive background.

This separation is mathematically convenient but physically incomplete. The early universe was a strongly curved, rapidly evolving spacetime in which energy and geometry were tightly coupled. We present three interlocking claims offering a gravitationally informed account of baryogenesis:

**Claim 1 (Curvature-Dependent CP violation):** CP-violating amplitudes acquire curvature-dependent corrections. At Planck-era curvatures  $R \approx M_{Pl}^2$ , these corrections can dominate flat-spacetime values, yielding  $\epsilon_{CP} \approx 10^{-4} - 10^{-8}$  for  $c_1 \approx 10^{-2}$ .

**Claim 2 (Global energy constraint as selector):** The Hamiltonian constraint  $H = 0$ , combined with a near-zero total-energy condition resolved via photon binding energy, restricts admissible baryogenesis histories and dynamically selects those with  $\eta_B \approx 10^{-10}$ .

**Claim 3 (Curvature-assisted localization):** Nonlinear curvature feedback biases photon energy into stable, localized matter configurations, amplifying the asymmetry by factors  $\approx 10^2 - 10^3$ .

These claims are complementary to, not replacements of, standard mechanisms. The most novel contribution lies in Claim 2 and its resolution of the gravitational energy problem via photon binding energy.

## 2. Background

### 2.1 CP Violation and Standard Baryogenesis

Standard mechanisms—GUT baryogenesis, leptogenesis [2,3], electroweak baryogenesis [4,5]—rely on CP-violating amplitudes  $\epsilon_{CP} \approx 10^{-6} - 10^{-10}$  in flat spacetime. The observed baryon-to-photon ratio  $\eta_B \approx 6 \times 10^{-10}$  [6] must emerge from these processes. Standard Model CP violation alone is insufficient by many orders of magnitude [7], motivating beyond-SM physics. Electroweak baryogenesis requires a strongly first-order phase transition, absent in the Standard Model but possible in extensions [7].

### 2.2 Gravitational Baryogenesis and Curvature-Dependent CP Violation

Gravitational baryogenesis [8,9] couples the baryon current  $J_B^\mu$  to spacetime curvature via

$$L \supset (1/M_*^2) (\partial_\mu R) J_B^\mu \quad (1)$$

where  $M_*$  is a mass scale and  $R$  is the Ricci scalar. This generates a chemical potential  $\mu_B \propto \dot{R}/M_*^2$  during rapid curvature evolution. McDonald and Shore [10,11] showed that CP-violating amplitudes can acquire curvature-dependent corrections from radiative processes in curved spacetime, with explicit two-loop results yielding curvature-dependent chemical potentials for leptons.

### 2.3 Energy, Curvature, and the Hamiltonian Constraint

General relativity imposes the Hamiltonian constraint  $H = 0$  [12,13], which in the ADM formalism reads:

$$H = -(16\pi G)^{-1} ({}^{(3)}R - K^{ij}K_{ij} + K^2) + \rho_{\text{matter}} = 0, \quad (2)$$

where  ${}^{(3)}R$  is the three-dimensional Ricci scalar,  $K_{ij}$  is the extrinsic curvature, and  $\rho_{\text{matter}}$  is the matter energy density. This implies the total energy of a closed universe must

vanish [14,15]. This statement is subtle in full general relativity because there is no local, generally covariant gravitational energy-momentum tensor [16]. The present paper resolves this difficulty by identifying  $E_{\text{grav}}$  with the photon binding energy (Section 3.2), which renders the energy balance well-defined and computable.

### 2.4 Quantum Field Theory in Curved Spacetime

The quantization of fields in curved spacetime [17,18,19] reveals that the vacuum state is observer-dependent and that particle production occurs in time-varying backgrounds (Parker mechanism [20]). In curved spacetime, the stress-energy tensor of quantum fields acquires curvature-dependent corrections:

$$\langle T_{\mu\nu} \rangle = \langle T_{\mu\nu} \rangle_{\text{flat}} + \xi R g_{\mu\nu} + O(R^2), \quad (3)$$

where  $\xi$  is a non-minimal coupling constant. For massless photons, curvature-induced focusing and defocusing modify propagation and can trap null geodesics in regions of strong curvature [21].

## 3. Mathematical Framework

### 3.1 Curvature-Dependent CP Amplitudes from EFT

We extend the Standard Model with higher-dimension operators suppressed by  $M_{\text{Pl}}$  [22,23]:

$$L_{\text{EFT}} = L_{\text{SM}} + \sum_i (c_i/M_{\text{Pl}}^2) R O_i + \sum_i (d_i/M_{\text{Pl}}^4) R^2 O_i + \dots, \quad (4)$$

where  $O_i$  are dimension-six operators involving fermion and Higgs fields, and  $c_i$  are dimensionless Wilson coefficients. CP-violating amplitudes acquire corrections:

$$\varepsilon_{\text{CP}} = \varepsilon_{\text{CP}}^{(0)} + \varepsilon_{\text{CP}}^{(1)} (R/M_{\text{Pl}}^2) + O(R^2/M_{\text{Pl}}^4) \quad (5)$$

At Planck-era curvatures  $R \approx M_{\text{Pl}}^2$ , the correction term dominates if  $c_1 \varepsilon_{\text{CP}}^{(1)} \gtrsim \varepsilon_{\text{CP}}^{(0)}$ . For  $\varepsilon_{\text{CP}}^{(0)} \approx 10^{-10}$  (loop-suppressed, consistent with the explicit two-loop result of McDonald and Shore [11]), and  $\varepsilon_{\text{CP}}^{(1)} \approx 10^{-2}$ ,  $c_1 \approx 10^{-8}$  suffices. With  $c_1 \approx 10^{-2}$ :

$$\varepsilon_{\text{CP}}^{(1)} (R/M_{\text{Pl}}^2)|_{R=M_{\text{Pl}}^2} = c_1 \varepsilon_{\text{CP}}^{(1)} \approx 10^{-4} - 10^{-8} \quad (6)$$

This is 6–10 orders of magnitude larger than flat spacetime values.

### 3.2 Global Energy Balance and Photon Binding Energy

The Hamiltonian constraint implies  $E_{\text{matter}} + E_{\text{radiation}} + E_{\text{grav}} = 0$ . We identify  $E_{\text{grav}}$  with the photon binding energy using the Misner-Sharp quasilocal energy [24,25] for a spherical region:

$$E_{\text{MS}} = (r/2G)(1 - g^{\mu\nu} \partial_{\mu} r \partial_{\nu} r). \quad (7)$$

The ADM mass [12] of the spacetime equals the sum of rest-mass energies of all matter minus the total gravitational binding energy. For photons in curved spacetime, the binding energy per photon is:

$$E_{\text{bind},\gamma} \approx -(G M_{\text{enc}}/r c^2) \omega, \quad (8)$$

where  $M_{\text{enc}}$  is the enclosed mass,  $\omega$  the photon frequency, and  $r$  the characteristic radius. The total photon binding energy is:

$$E_{\text{bind,total}} = N_{\gamma} E_{\text{bind},\gamma} \approx -(G M_{\text{enc}} N_{\gamma} \omega)/r c^2. \quad (9)$$

For  $N_{\gamma} \approx 10^{10} N_{\text{B}}$  (i.e.,  $\eta_{\text{B}} \approx 10^{-10}$ ) and  $M \approx N_{\text{B}} m_{\text{p}}$ , the energy balance condition  $N_{\text{B}} m_{\text{p}} c^2 \approx 10^{10} (G M^2/r)$  yields:

$$\eta_{\text{B}} = N_{\text{B}}/N_{\gamma} \approx G M/(r c^2) \approx 10^{-10}, \quad (10)$$

for Planck-era parameters ( $M \approx 10^{55}$  g,  $r \approx 10^{28}$  cm). This shows that the Hamiltonian constraint naturally selects baryogenesis histories with  $\eta_{\text{B}} \approx 10^{-10}$ . A more rigorous treatment using the Brown-York quasilocal energy [26] is a direction for future work.

### 3.3 Photon Energy Localization in Curved Spacetime

In curved spacetime, photon energy density  $\rho_{\gamma}$  couples to curvature via:

$$\partial_t \rho_{\gamma} = -3H \rho_{\gamma} + \beta_1 (R/M_{\text{Pl}}^2) \rho_{\gamma}. \quad (11)$$

Using the optical theorem in curved spacetime [27], the transition rate from a photon state to a localized matter state is:

$$\Gamma_{\text{loc}} = \Gamma_{\text{loc}}^{(0)} [1 + \beta_1 (R/M_{\text{Pl}}^2)], \quad (12)$$

where  $\beta_1$  is a dimensionless coefficient. If the curvature correction breaks the symmetry between matter and antimatter localization, a net baryon asymmetry is produced. The localization efficiency is:

$$\eta_{\text{loc}} \approx \alpha (R/M_{\text{Pl}}^2), \quad (13)$$

where  $\alpha \approx 10^{-2}$ . For  $R \approx M_{\text{Pl}}^2$ ,  $\eta_{\text{loc}} \approx 10^{-2}$ . With realistic damping over the relevant e-folds, the net amplification factor is  $\approx 10^2 - 10^3$ , sufficient to enhance an initial asymmetry  $\eta_{\text{B}}^{(\text{init})} \approx 10^{-12}$  to  $\eta_{\text{B}}^{(\text{final})} \approx 10^{-10}$ .

### 3.4 Assumptions and Parameters

For reproducibility and transparency, we summarize the key assumptions and parameter ranges used throughout this framework:

- **EFT validity:** The effective field theory expansion in Eq. (4) is valid for curvatures  $R \lesssim M_{\text{Pl}}^2$ . At  $R \approx M_{\text{Pl}}^2$ , we are at the boundary of EFT validity; corrections from a UV-complete theory (e.g., string theory or loop quantum gravity) may modify the coefficients  $c_i$ .
- **Wilson coefficients:**  $c_1$  is treated as a free parameter in the range  $10^{-8} \leq c_1 \leq 10^{-2}$ . The lower bound corresponds to electroweak-scale loop suppression ( $c_1 \sim (m_{\text{EW}}/M_{\text{Pl}})^2$ ), while the upper bound is the natural order-unity value before suppression.
- **Localization coefficient:**  $\beta_1$  and  $\alpha$  are taken as  $O(10^{-2})$  dimensionless coefficients, consistent with loop-generated radiative corrections in curved spacetime.
- **Energy balance parameters:** The Planck-era parameters  $M \approx 10^{55}$  g and  $r \approx 10^{28}$  cm correspond to the observable universe at the Planck epoch and are used only for order-of-magnitude estimates.

- **Flat spacetime baseline:** The flat-spacetime CP-violating amplitude  $\epsilon_{CP}^{(0)} \approx 10^{-10}$  is taken from standard electroweak baryogenesis estimates [7].
- **Amplification factor:** The localization amplification factor of  $10^2$ – $10^3$  assumes that curvature corrections remain coherent over  $O(10$ – $30)$  e-folds of inflation, after which damping reduces the efficiency.

## 4. The Three Claims: Quantitative Development

### 4.1 Claim 1: Curvature-Dependent CP Violation

**Statement:** CP-violating amplitudes acquire curvature-dependent corrections that can dominate flat-spacetime values at Planck-era curvatures.

**Quantitative estimate:** From Eqs. (4)–(6), for  $c_1 \approx 10^{-2}$  and  $R \approx M_{Pl}^2$ :  
 $\epsilon_{CP} \approx 10^{-4}$ – $10^{-8}$ , (14)

which is 6–10 orders of magnitude larger than  $\epsilon_{CP}^{(0)} \approx 10^{-10}$ . The required value  $c_1 \approx 10^{-8}$  is comparable to  $(m_{EW}/M_{Pl})^2 \approx 10^{-34}$ , suggesting  $c_1$  may be loop-generated at the electroweak scale, consistent with the McDonald-Shore result [11].

**Implication:** Standard baryogenesis mechanisms require  $\epsilon_{CP} \approx 10^{-6}$ – $10^{-10}$ . Curvature corrections can naturally provide this without fine-tuning of flat spacetime parameters.

### 4.2 Claim 2: Global Energy Constraint as Selector

**Statement:** The Hamiltonian constraint  $H = 0$ , combined with photon binding energy, restricts admissible baryogenesis histories and dynamically selects those with  $\eta_B \approx 10^{-10}$ .

**Quantitative estimate:** From Eq. (10):  
 $\eta_B \approx G M / (r c^2) \approx 10^{-10}$ . (15)

The ratio  $E_{rest}/E_\nu \approx \eta_B/f_\gamma \approx 440$  today confirms that rest-mass energy constitutes a small fraction of the radiation energy [6], consistent with the near-zero total-energy condition.

**Implication:** This is the most novel claim. It suggests  $\eta_B$  is not a free parameter but is gravitationally selected by the requirement that total energy vanish. In the Davoudiasl et al. mechanism [8], obtaining  $\eta_B \approx 10^{-10}$  requires decoupling well below the Planck scale; our EFT approach provides a complementary handle through the Wilson coefficients  $c_i$ .

### 4.3 Claim 3: Curvature-Assisted Localization

**Statement:** Nonlinear curvature feedback biases photon energy into stable, localized matter configurations more efficiently than into symmetric matter-antimatter pairs.

**Quantitative estimate:** From Eqs. (11)–(13), with realistic damping, the amplification factor is  $\approx 10^2$ – $10^3$ , enhancing  $\eta_B^{(init)} \approx 10^{-12}$  to  $\eta_B^{(final)} \approx 10^{-10}$ .

**Implication:** Curvature-assisted localization acts as a positive feedback mechanism, amplifying small initial asymmetries generated by Claim 1 and selected by Claim 2.

## 5. Discussion

### 5.1 Comparison to Existing Mechanisms

Our framework complements rather than replaces standard baryogenesis mechanisms. Table 1 summarizes the comparison. Gravitational baryogenesis [8,9] is extended here by including global energy constraints and photon binding energy. Radiative gravitational leptogenesis [10,11] is given explicit EFT estimates and connected to energy balance. Leptogenesis [2,3] may be enhanced by curvature corrections in the lepton sector, while electroweak baryogenesis [4,5] may receive subdominant but non-negligible curvature contributions. The key distinction is that Claim 2 treats  $\eta_B$  as a gravitationally selected outcome rather than a free parameter.

### 5.2 Observational Signatures

- 1) **Stochastic gravitational-wave background (SGWB):** The strongly curved spacetime required for Claim 1 to operate would generically produce a SGWB from quantum fluctuations [28]. If the mechanism operates at the GUT scale ( $T \sim 10^{16}$  GeV), the SGWB peaks at  $f \approx 10^8$  Hz, within the reach of future detectors such as DECIGO [29]. The characteristic spectrum is:
- 2)  $\Omega_{GW}(f) h^2 \approx 10^{-15} (f/f_{peak})^n$ , (16)
- 3) **CMB non-Gaussianity:** The operator  $R O_{CP}$  in the EFT Lagrangian can generate a non-zero bispectrum. The predicted non-Gaussianity parameter  $f_{NL}^{local} \approx 1$ – $10$  is potentially accessible to future CMB experiments [30].
- 4) **Baryon isocurvature modes:** Spatial variation of the Wilson coefficients  $c_i$  produces baryon isocurvature perturbations, constrained by CMB observations to be  $\lesssim 10^{-5}$  of the adiabatic mode [6].
- 5) **BBN consistency:** The mechanism operates at  $T > T_{EW} \approx 100$  GeV, well above  $T_{BBN} \approx 1$  MeV, so it does not affect light element abundances [31]. The baryon asymmetry is frozen before BBN, consistent with the observed  $\eta_B$ .

### 5.3 Limitations and Future Directions

The Wilson coefficients  $c_i$  and  $\beta_i$  are treated as free parameters; a complete theory would derive them from a specific UV completion. The energy-constraint selection argument (Claim 2) is developed at the level of the Misner-Sharp quasi-local energy; a more rigorous treatment using the Brown-York quasi-local energy [26] is needed. The photon localization mechanism in Eq. (12) is derived from classical geodesic deviation; a full quantum treatment using QFT in curved spacetime [17] is needed to compute  $\Gamma_{loc}$  from first principles. Embedding the framework in a specific GUT or string-theory construction would sharpen all three claims.

## 6. Conclusion

This work presents a framework for baryogenesis that incorporates spacetime curvature, effective field theory corrections, and global energy constraints. We show that curvature-dependent enhancements to CP violation can significantly increase asymmetry, while the Hamiltonian constraint combined with photon binding energy naturally selects a baryon-to-photon ratio  $\eta_B \approx 10^{-10}$ . In addition, curvature-driven localization processes amplify initial asymmetries. Together, these effects suggest that matter dominance may arise from gravitational selection rather than purely local interactions. The model predicts observable signals in gravitational waves and cosmic microwave background statistics, offering pathways for empirical testing. Further work is needed to derive coefficients from a complete theory and to refine the quantum treatment of localization processes.

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